

Observational Radiative Kernels and Contributions to TOA Flux Anomalies

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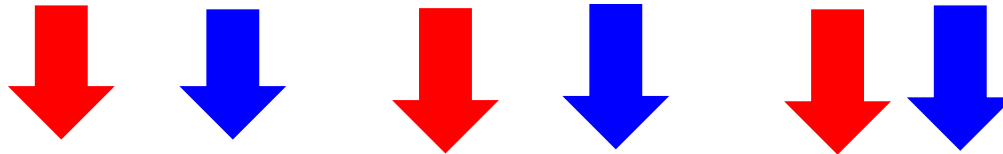
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October 2, 2018

NASA Sounders Meeting, 2018

Feedbacks and Radiative Kernels

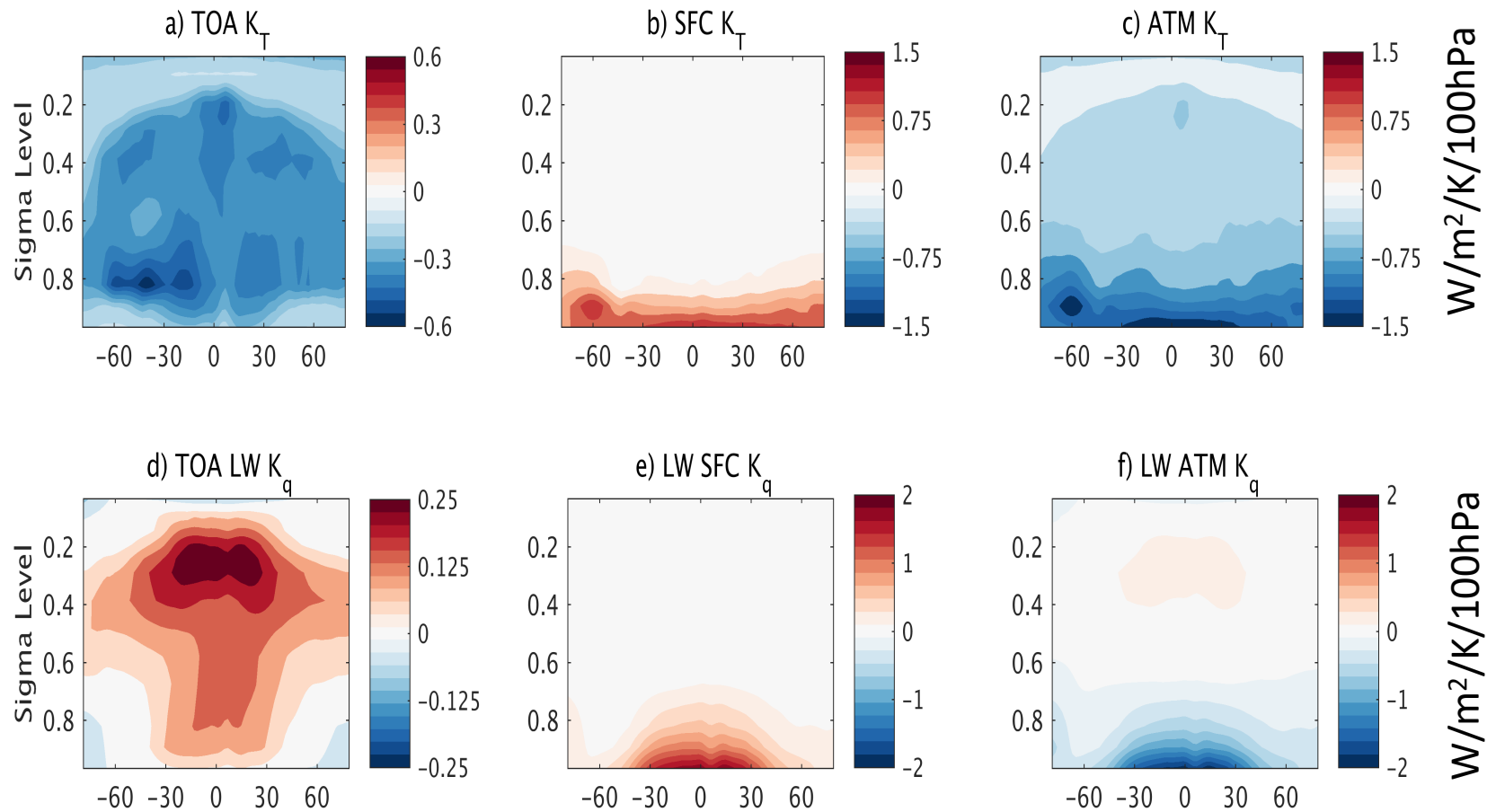


$$\lambda = -\frac{dR}{dT_s} = \frac{\delta R}{\delta T} \frac{dT}{dT_s} + \frac{\delta R}{\delta W} \frac{dW}{dT_s} + \frac{\delta R}{\delta \alpha} \frac{d\alpha}{dT_s} + \frac{\delta R}{\delta C} \frac{dC}{dT_s}$$

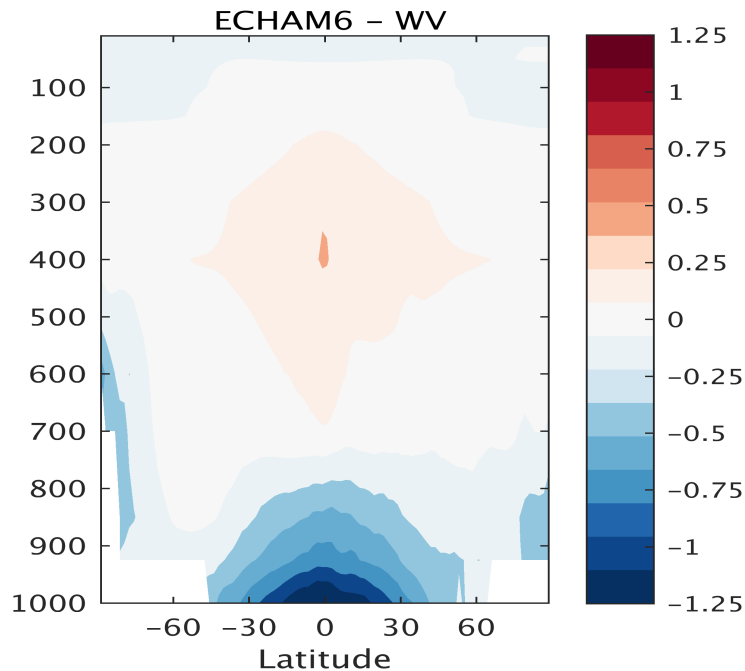
Feedback = Kernels X Climate Response

$$\frac{dR_c}{dT_s} = \frac{dCRF}{dT_s} + \sum (K_0^x - K^x) dx / dT_s$$

Radiative Kernels

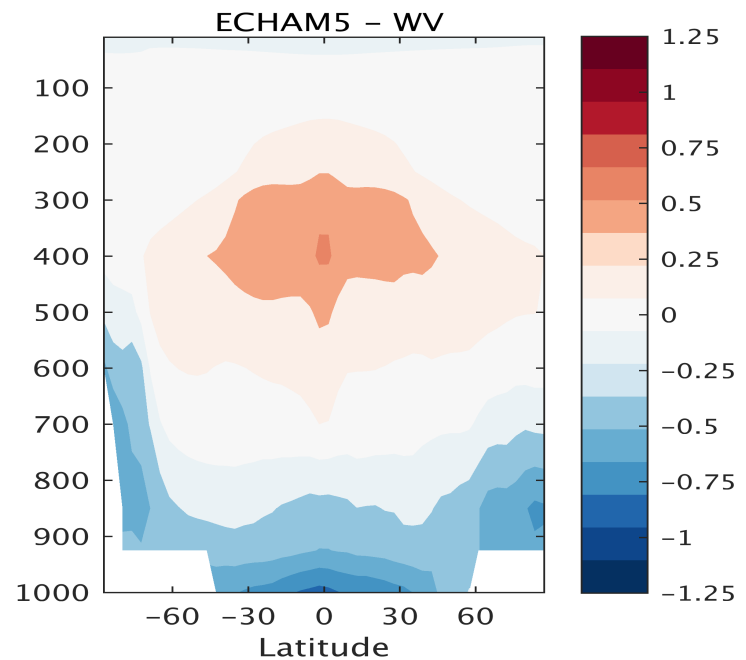


Differing spatial distribution of LW Water Vapor kernel



$$\lambda_{WV,LW} = -0.66 \text{ W/m}^2/\text{K}$$

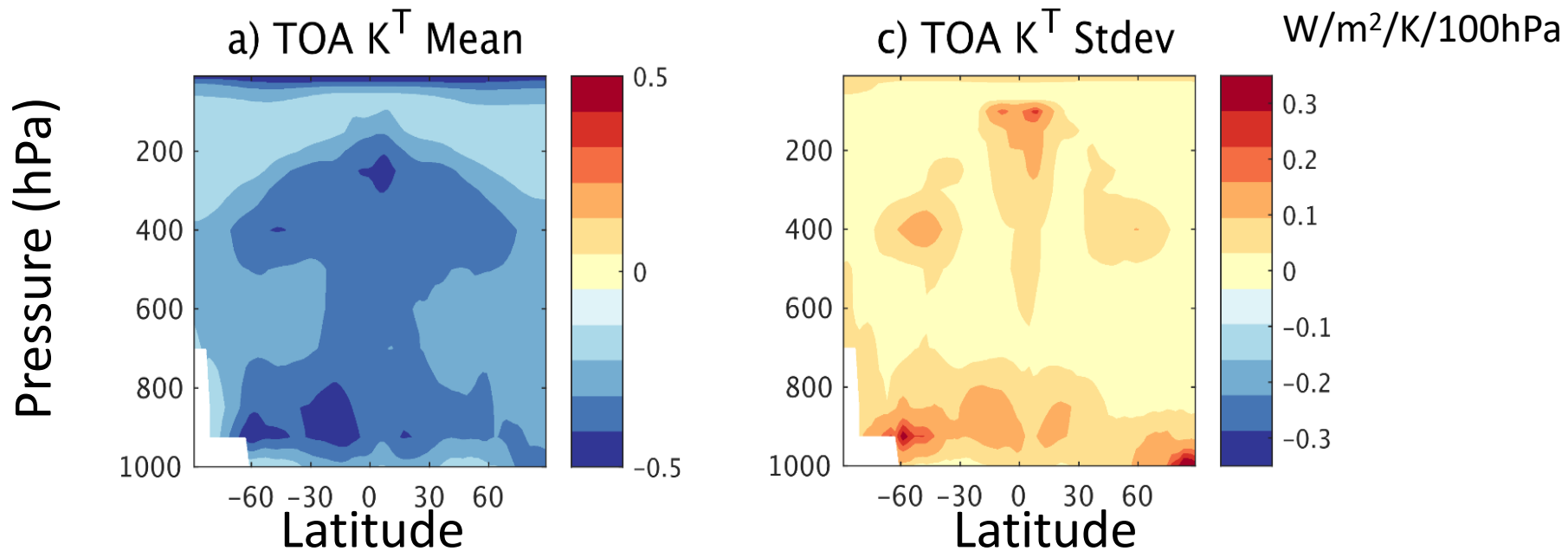
Flaschner et al. 2016



$$\lambda_{WV,LW} = 0.29 \text{ W/m}^2/\text{K}$$

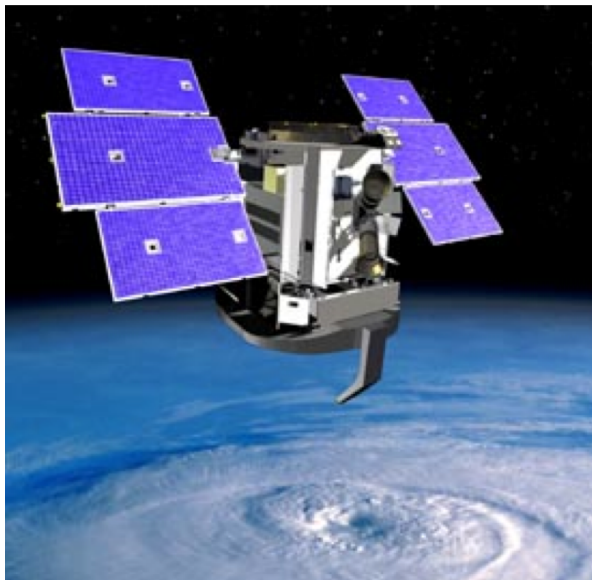
O'Gorman et al. 2012

Contributions to Radiative Kernel Differences



Kernel sensitivity to cloud climatology

CloudSat Radiative Kernels

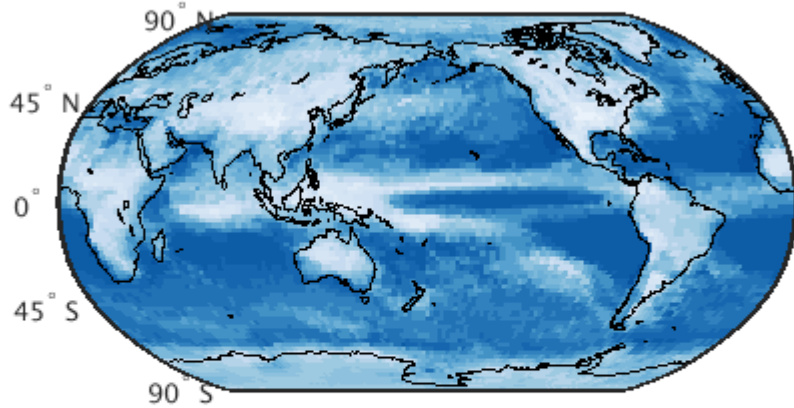


Observational Kernels from 2B-FLXHR-LIDAR:

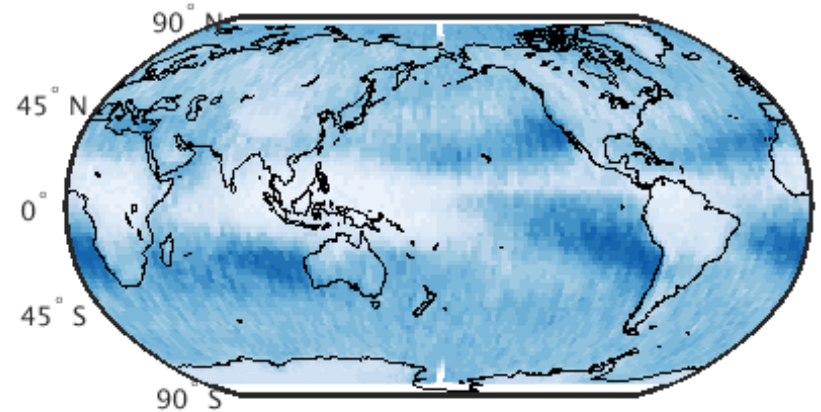
Input	Source
Cloud Properties	CloudSat, CALIPSO
Temperature, Humidity	ECMWF Reanalysis
Surface Albedo	MODIS, AMSR-E
Aerosol Properties	CALIPSO

Impact of Cloud Field on TOA T Kernels

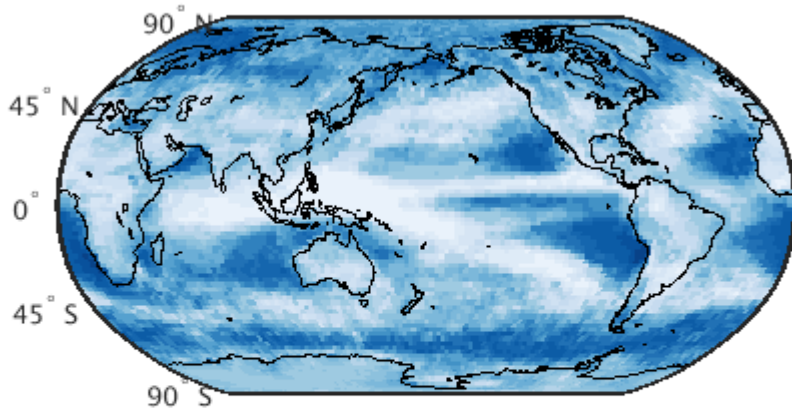
CESM Max K^T Pressure



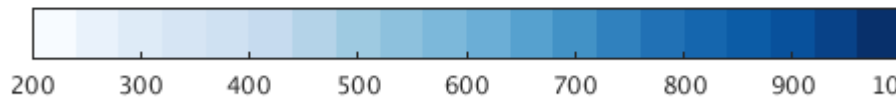
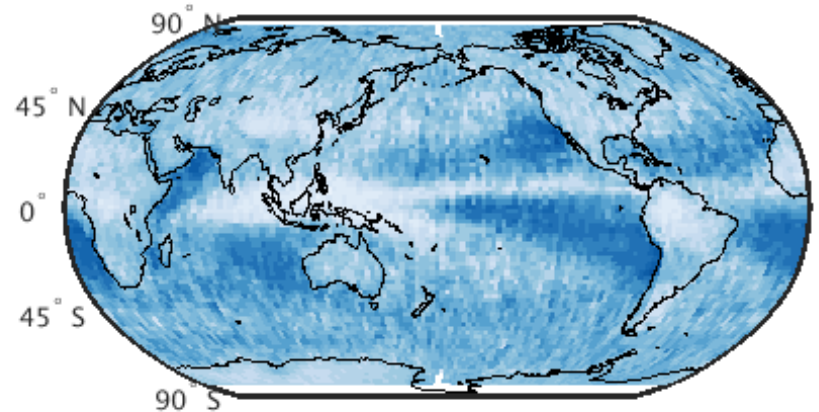
CloudSat Cloud Top Pressure



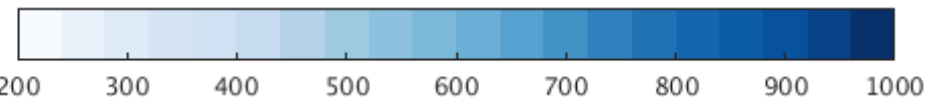
GFDL Max K^T Pressure



CloudSat Max K^T Pressure



hPa

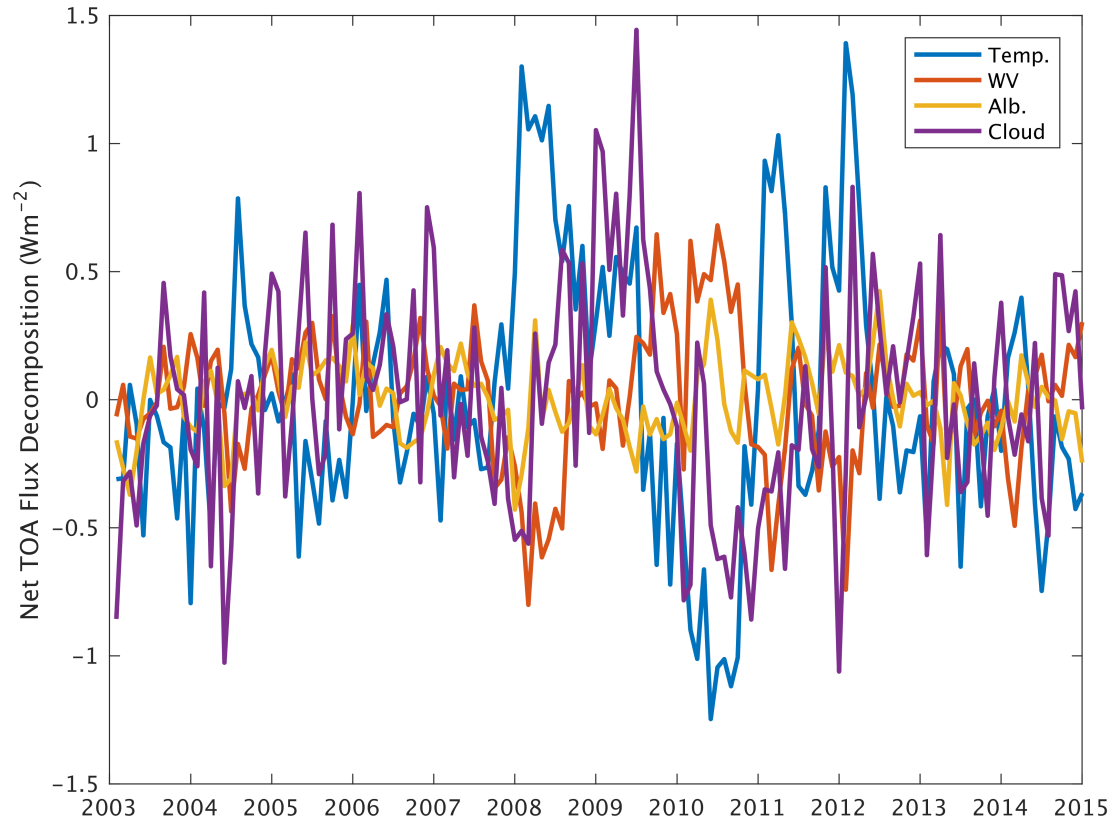


Kramer et al. 2018a

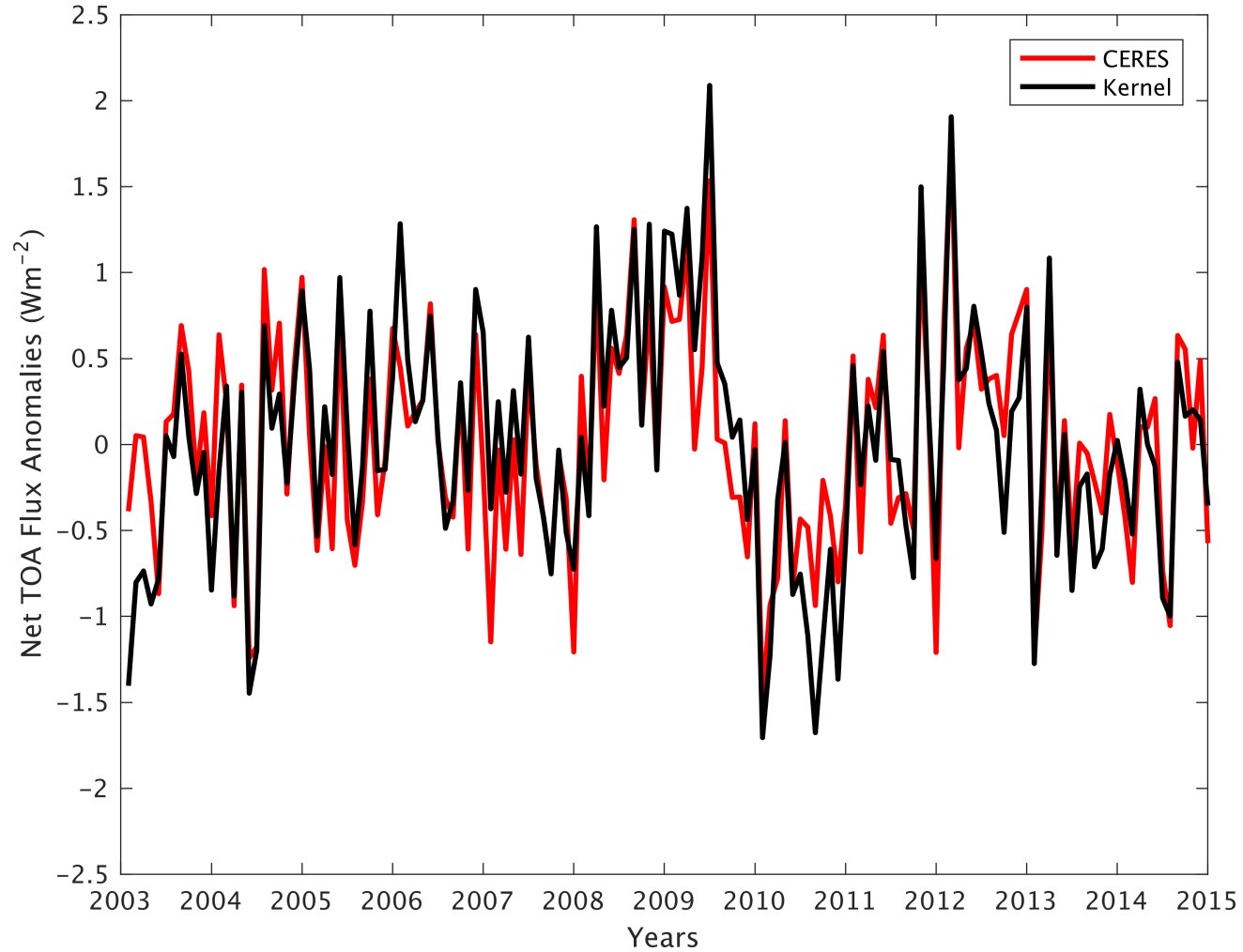
Individual Contributions to TOA Flux Anomalies using AIRS

$$\Delta R_{\text{TOA}} = \sum \Delta R_x + F$$

$$\Delta R_x = K_x * dx$$

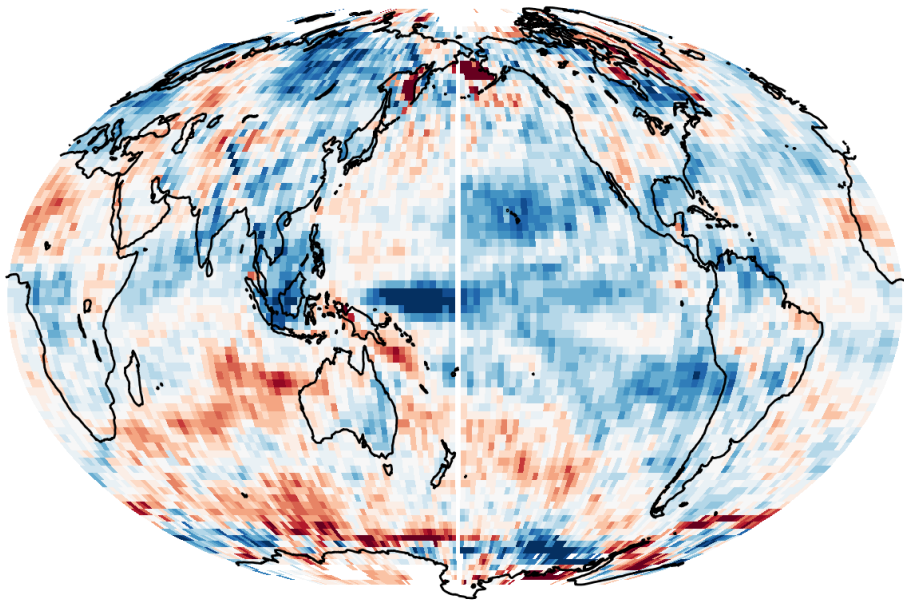


Net TOA Flux Anomalies

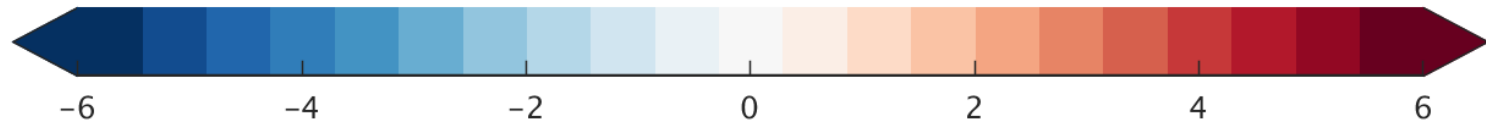
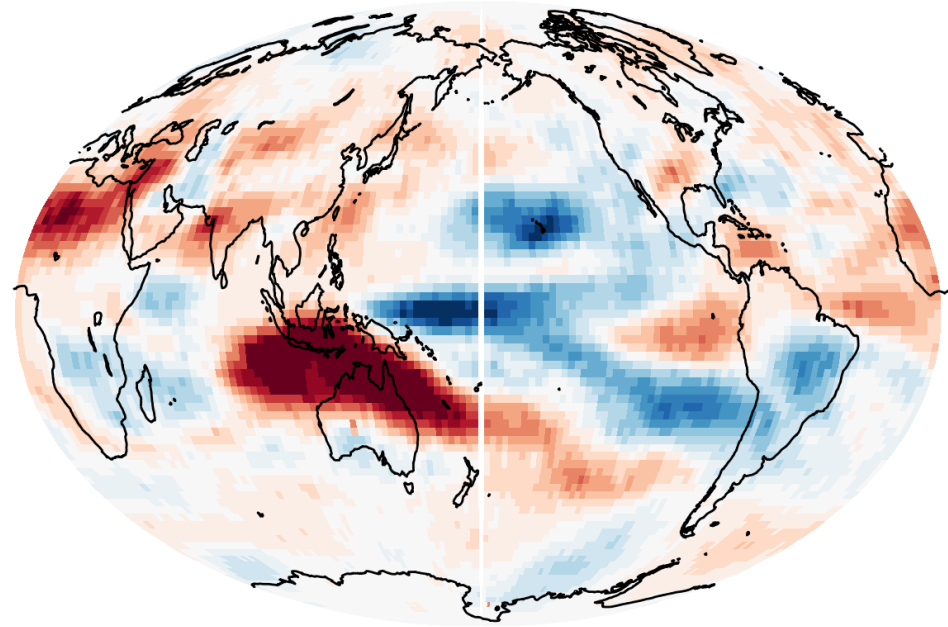


Flux Anomalies 6/2010-10/2010

Net Δ TOA (CERES-Kernel)



Water Vapor Flux

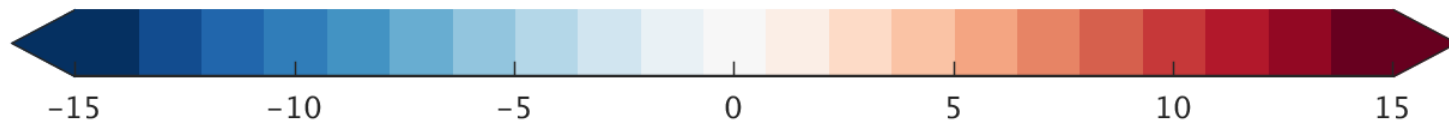
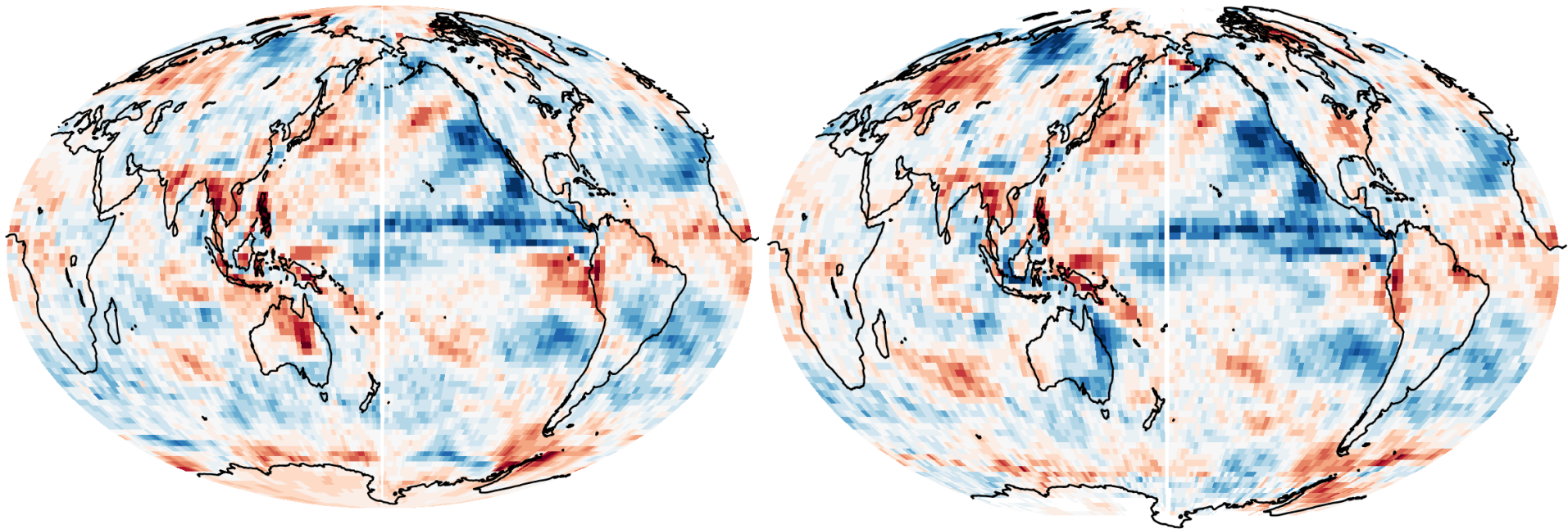


W/m²

Flux Anomalies 6/2010-10/2010

Net TOA CERES

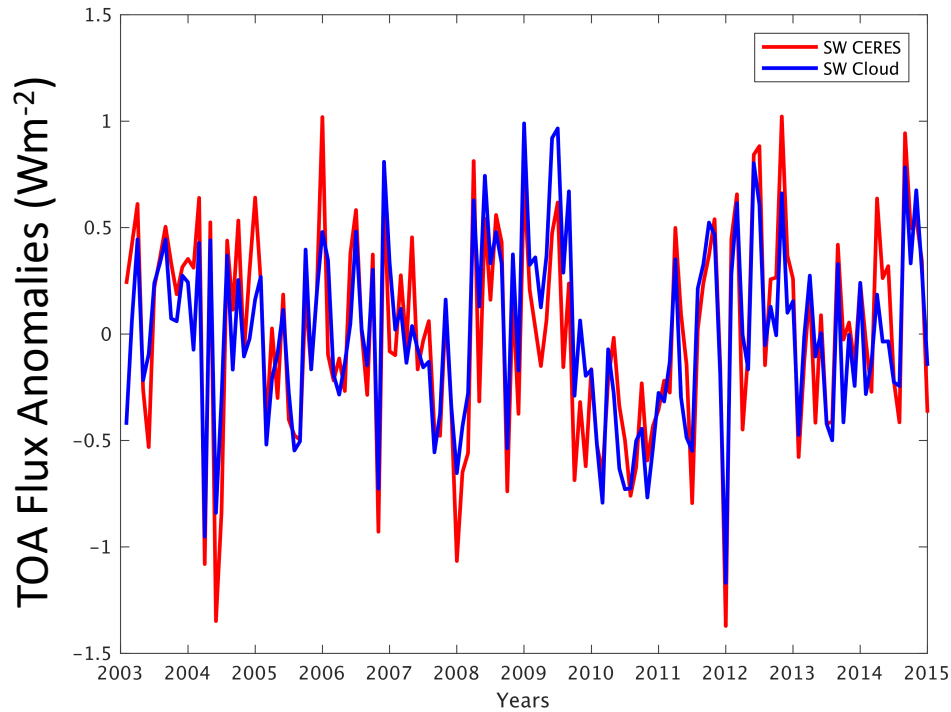
Cloud Fluxes



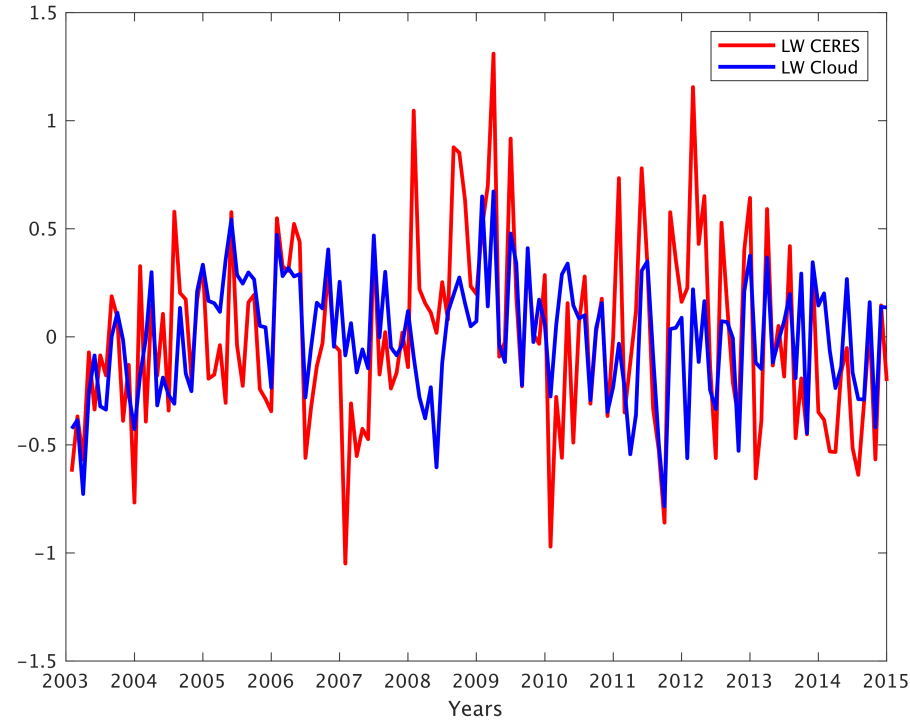
W/m^2

Cloud Flux Contribution

SW Cloud Flux

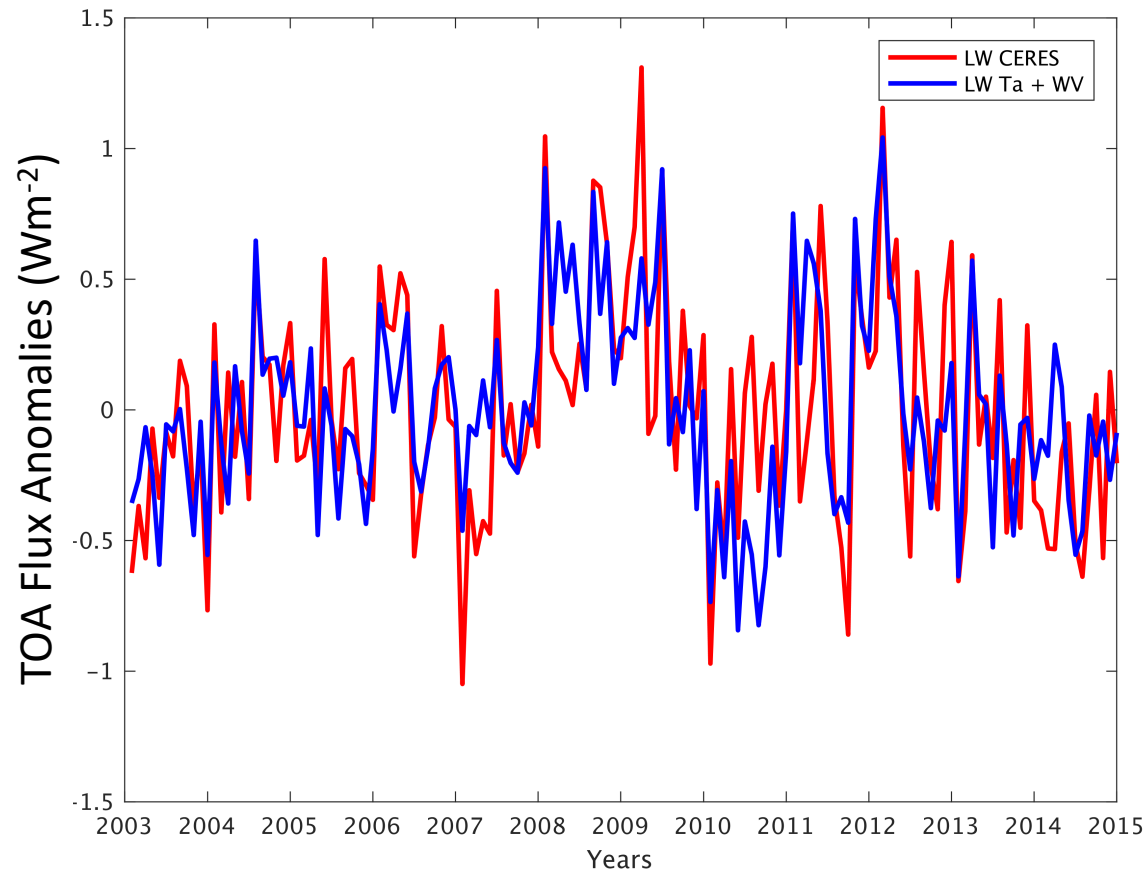


LW Cloud Flux



Non-Cloud Flux Contribution

LW Ta +WV Flux

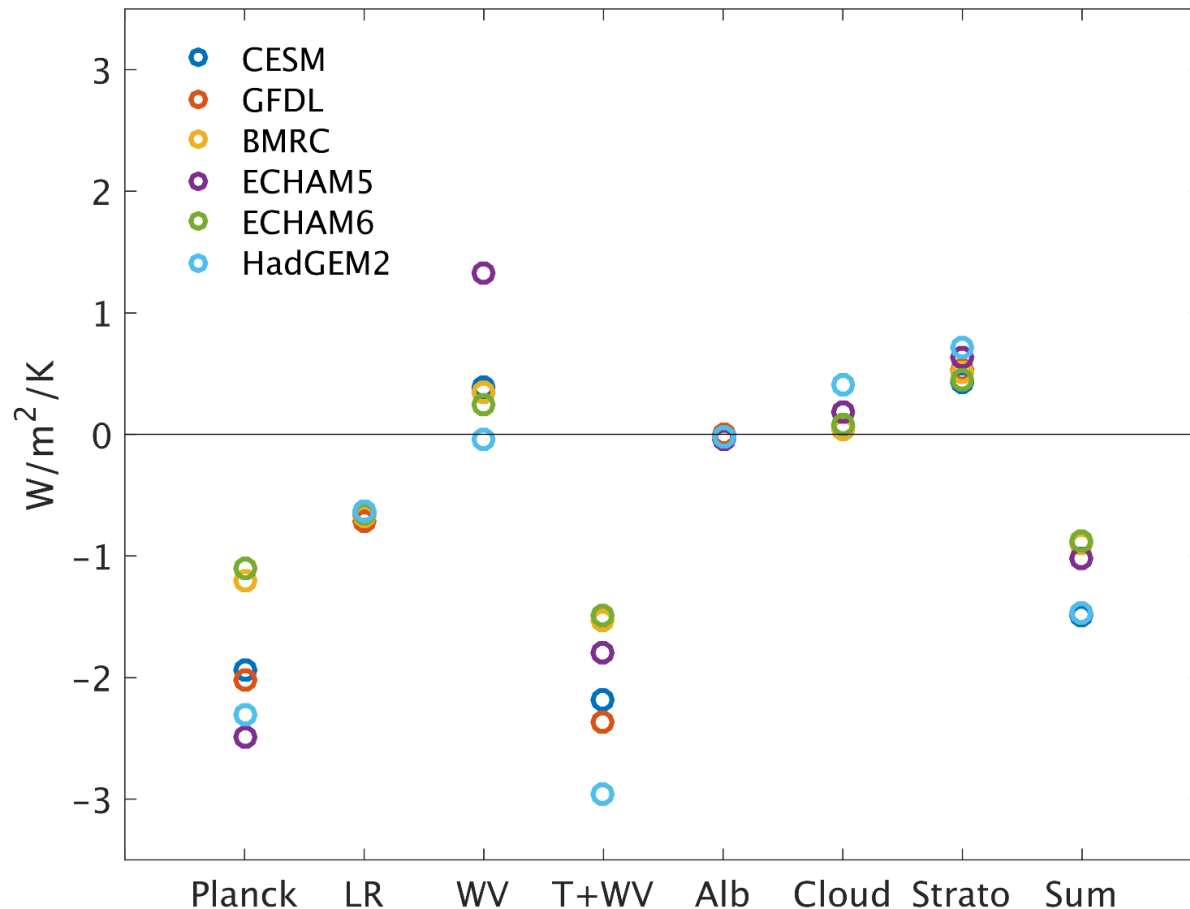


Summary

- Documented differences in atmospheric GCM-based radiative kernels
- GCM-based Radiative kernels are subject to biases in climatological clouds. Instead we can use CloudSat-based radiative kernels.
- Clouds explain variability in observed SW TOA flux anomalies while non-cloud changes explain LW TOA

Estimating Feedbacks with different Radiative Kernels

CMIP5 Model-Mean ATM Feedbacks – Kernel Comparison



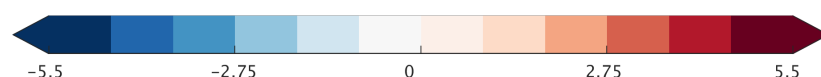
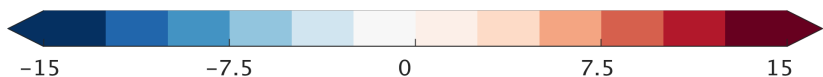
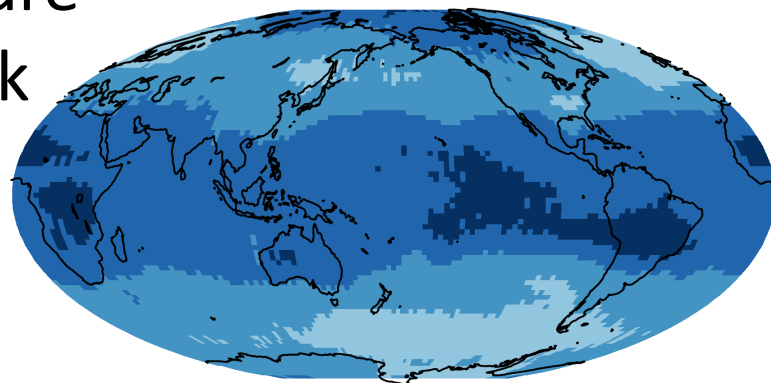
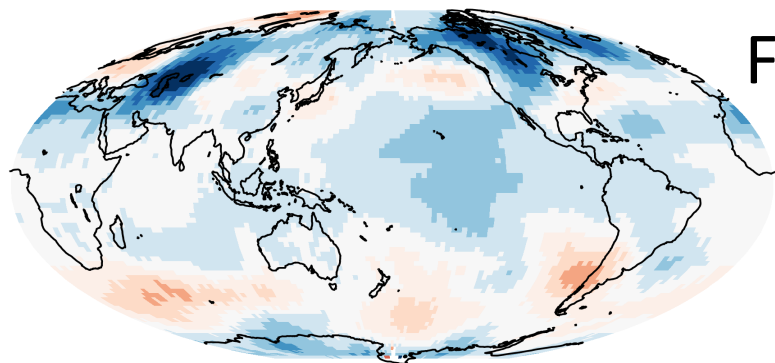
- ECHAM5 WV Kernel is an outlier
- Planck spread associated with near-surface vertical resolution differences and RTM artifacts
- Stratospheric response sensitive to vertical resolution near TOA
- Cloud feedback spread due to cloud masking differences

Inter-annual versus Climate Change Feedbacks

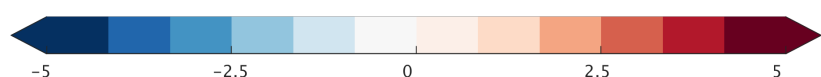
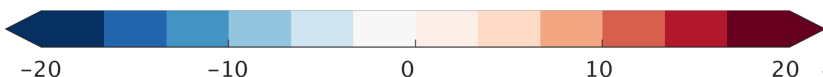
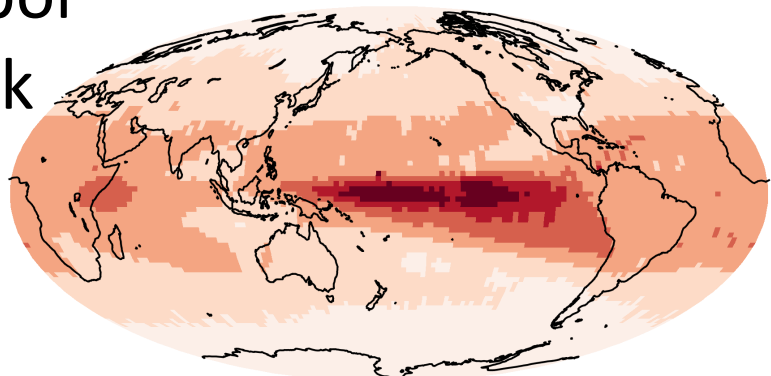
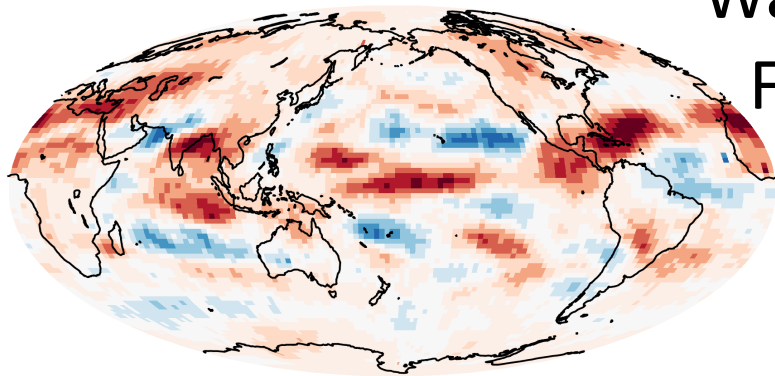
Observed

Temperature
Feedback

Long-term (CMIP5)



Water Vapor
Feedback



W/m²/K

Feedback magnitude consistent between timescales

